#### **SET 4 – GRAVITATION**

(a)  $6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2$ 

(b) 9.8 N/kg

1. The universal law of gravitation was proposed by:  (a) Newton  (b) Cavendish  (c) Kepler  (d) Einstein
2. Gravitational force between two masses varies directly with:  (a) product of their masses  (b) sum of their masses  (c) difference of their masses  (d) square of their masses
3. Gravitational force between two masses varies inversely with:  (a) square of distance  (b) distance  (c) cube of distance  (d) square of their product
<b>4.</b> The dimensional formula of G is:  (a) [M <sup>-1</sup> L <sup>3</sup> T <sup>-2</sup> ]  (b) [ML <sup>2</sup> T <sup>-2</sup> ]  (c) [MLT <sup>-2</sup> ]  (d) [M <sup>-2</sup> L <sup>3</sup> T <sup>-2</sup> ]
5. SI unit of G is:  (a) N·m²/kg²  (b) N/kg²  (c) m²/kg²  (d) N·m/kg
6. The value of G is:

(c) 3 × 10 <sup>8</sup> m/s (d) 1.6 × 10 <sup>-19</sup> C	
<b>7.</b> The value of G is:	
(a) constant everywhere	
(b) depends on location	
(c) depends on medium	
(d) depends on mass	
8. Gravitational force is:	
(a) always attractive	
(b) always repulsive	
(c) zero	
(d) both attractive and repulsive	
<ul><li>9. The value of acceleration due to gravity at (a) maximum</li><li>(b) minimum</li><li>(c) zero</li><li>(d) equal to that at equator</li></ul>	at poles is:
10. The value of acceleration due to gravity	at Earth's centre is:
(a) zero	
(b) maximum	
(c) infinite	
(d) 9.8 m/s <sup>2</sup>	
<b>11.</b> The value of g on Moon is about:	
(a) 1/6th of Earth's	
(b) 6 times Earth's	
(c) 1/3rd of Earth's	
(d) equal to Earth's	
<b>12.</b> Gravitational field intensity is defined as	:

(a) force per unit mass(b) force per unit volume

<ul><li>(c) mass per unit force</li><li>(d) acceleration per unit time</li></ul>	
<b>13.</b> Gravitational field intensity near Earth's surface in	s:
(a) 9.8 N/kg	
(b) 9.8 m/s	
(c) 1 N/kg	
(d) 10 kg/N	
<b>14.</b> The potential at infinity is taken as:	
(a) zero	
(b) minimum	
(c) maximum	
(d) negative	
<b>15.</b> Gravitational potential energy between two bodies	es is:
(a) −Gm₁m₂/r	
(b) Gm₁m₂r	
(c) Gm <sub>1</sub> m <sub>2</sub> /r <sup>2</sup>	
(d) −Gm₁m₂r²	
<b>16.</b> Kepler's first law is the law of:	
(a) elliptical orbits	
(b) areas	
(c) gravitation	
(d) attraction	
17. Kepler's second law states that:	
(a) Equal areas are swept in equal time intervals	
(b) T <sup>2</sup> ∝ R <sup>3</sup>	
(c) $F \propto 1/r^2$	
(d) F = ma	
40 Kanlada third law since relative to the	
<b>18.</b> Kepler's third law gives relation between:	

(a) T<sup>2</sup> and R<sup>3</sup> (b) T and R<sup>2</sup>

(c) F and M (d) V and M
19. The gravitational potential energy is always:
(a) negative
(b) positive
(c) zero
(d) infinite
<b>20.</b> For a satellite of mass m at height h, potential energy is:
(a) −GMm/(R+h)
(b) GMm/(R+h)
(c) GMmR/h
(d) −GMmR/h²
21. The escape velocity is given by:
(a) √(2GM/R)
(b) √(GM/R)
(c) √(GM/2R)
(d) √(GR/M)
22. The escape velocity from Earth's surface is approximately:
(a) 11.2 km/s
(b) 9.8 km/s
(c) 7.9 km/s
(d) 8.2 km/s
23. The escape velocity depends on:
(a) radius and mass of the planet
(b) mass of object
(c) shape of object
(d) atmosphere
24. The ratio of escape velocity to orbital velocity is:
(a) √2

(b) 2

(c) 1 (d) ½	
25. The total energy of a satellite in circular orbit is:  (a) −GMm/2R	
(b) GMm/R	
(c) GMm/2R	
(d) -GMm/R	
26. The kinetic energy of a satellite in circular orbit is:	
(a) GMm/2R	
(b) -GMm/2R	
(c) GMm/R	
(d) −GMm/R	
27. The potential energy of a satellite in circular orbit is:  (a) -GMm/R  (b) GMm/R  (c) -GMm/2R  (d) GMm/2R	
28. The time period of a satellite is related to its orbital radius by:	
(a) $T^2 \propto R^3$	
(b) $T^2 \propto R^2$	
(c) $T \propto R^2$	
(d) $T^3 \propto R^2$	
29. The gravitational potential decreases with:	
(a) increase in distance	
<ul><li>(b) decrease in distance</li><li>(c) increase in mass</li></ul>	
(d) none	
(a) none	
30. The gravitational field is a:	
(a) conservative field	

(b) non-conservative field

(c) electric field (d) mechanical field	
<b>31.</b> A geostationary satellite revolves:	
(a) from west to east	
(b) from east to west	
(c) perpendicular to equator	
(d) through poles	
<b>32.</b> A geostationary satellite has a time period of:	
(a) 24 hours	
(b) 12 hours	
(c) 6 hours	
(d) 48 hours	
(a) 36,000 km (b) 3,600 km (c) 42,000 km (d) 64,000 km	
<b>34.</b> The orbital velocity of a satellite close to Earth	s:
(a) 7.9 km/s	
(b) 9.8 km/s	
(c) 10 km/s	
(d) 11.2 km/s	
35. The energy required to launch a satellite deper	ds on:
(a) mass and height of orbit	
(b) only mass	
(c) only height	
(d) none	
<b>36.</b> The gravitational potential at infinity is: (a) zero	
(a) 2010	

(b) positive

c) negative d) infinite
7. The weight of a body at the equator is:
a) less than at poles
b) greater than at poles
c) equal to poles d) zero
u) 2610
8. Weightlessness occurs when:
a) object is in free fall
b) object is at rest
c) object is stationary
d) g is maximum
9. The value of a decrease with:
<b>9.</b> The value of g decreases with: a) height, depth, and rotation
b) only height
c) only depth
d) only rotation
0. The gravitational field intensity at a distance r from mass M is:  a) GM/r² b) -GM/r² c) GM/r d) -GM/r³  1. The potential energy of a satellite is:  a) twice its total energy
b) equal to total energy
c) half of total energy
d) negative of total energy
2. Gravitational field intensity is: a) vector quantity

(b) scalar quantity

(c) tensor (d) none
43. Gravitational potential is:
(a) scalar quantity
(b) vector quantity
(c) tensor quantity
(d) none
<b>44.</b> The gravitational constant G is:
(a) same everywhere
(b) changes with height
(c) changes with mass
(d) depends on gravity
<b>45.</b> The value of g at height h << R is:
(a) g(1 – 2h/R)
(b) $g(1 - h/R)$
(c) $g(1 + h/R)$
(d) $g(1 + 2h/R)$
46. The escape velocity from the Moon is:
(a) 2.4 km/s
(b) 7.9 km/s
(c) 11.2 km/s
(d) 8.2 km/s
<b>47.</b> The gravitational potential energy is zero when:
(a) two bodies are infinitely apart
(b) two bodies touch
(c) at finite separation
(d) inside Earth
<b>48.</b> If distance between two bodies is tripled, gravitational force becomes:

(a) 1/9th(b) one-third

- (c) nine times
- (d) one-ninth
- **49.** The total mechanical energy of a satellite in circular orbit is:
- (a) negative
- (b) positive
- (c) zero
- (d) infinite
- **50.** Gravitational field lines never:
- (a) intersect
- (b) cross twice
- (c) start from infinity
- (d) bend

#### Answers – SET 4

- 1 (a) 2 (a) 3 (a) 4 (a) 5 (a) 6 (a) 7 (a) 8 (a) 9 (a) 10 (a)
- 11 (a) 12 (a) 13 (a) 14 (a) 15 (a) 16 (a) 17 (a) 18 (a) 19 (a) 20 (a)
- 21 (a) 22 (a) 23 (a) 24 (a) 25 (a) 26 (a) 27 (a) 28 (a) 29 (a) 30 (a)
- 31 (a) 32 (a) 33 (a) 34 (a) 35 (a) 36 (a) 37 (a) 38 (a) 39 (a) 40 (b)
- 41 (a) 42 (a) 43 (a) 44 (a) 45 (a) 46 (a) 47 (a) 48 (a) 49 (a) 50 (a)